

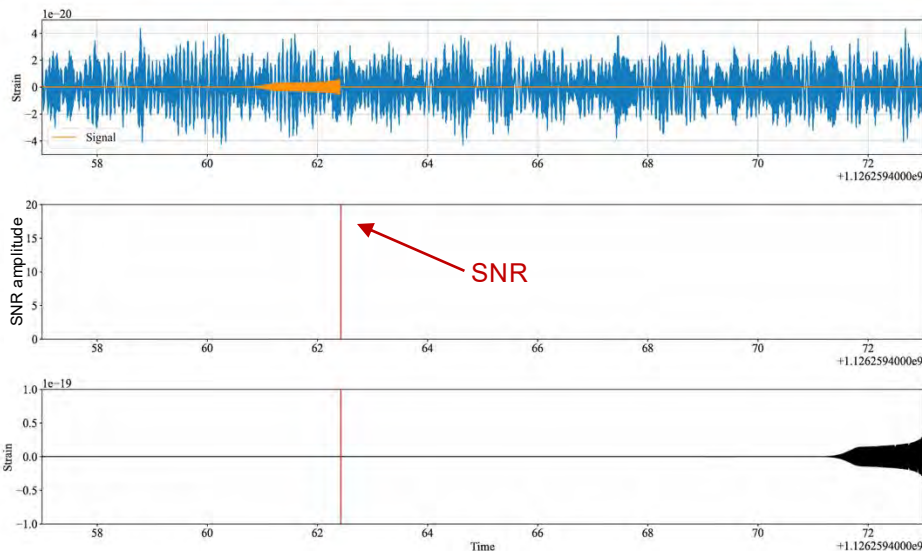


# Neural network classifiers for distinguishing signals from instrumental noise

Melissa Lopez  
ICERM 2025

# Modelled searches: matched filtering (MF) for CBC

What is matched filtering?



Unknown signal

\*

Template Bank  
(simulated GW)

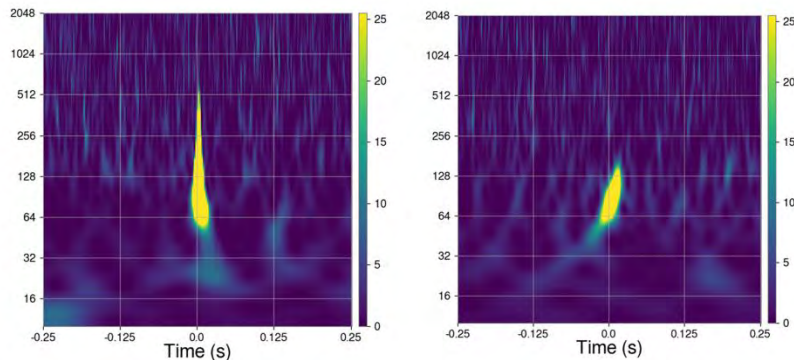
→

Triggers

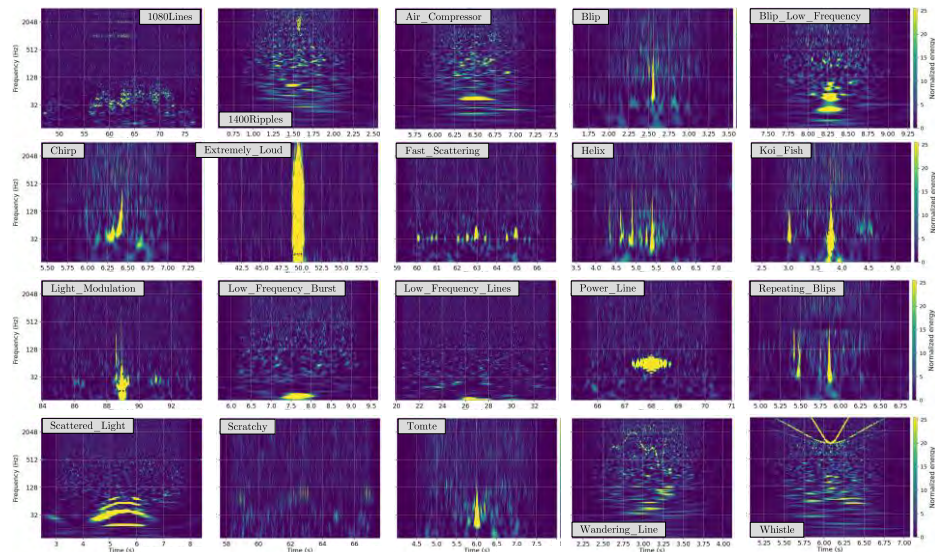
**Idea:** unknown signals generate *multiple* triggers. Can we find *patterns* with Machine Learning?

# Transient noise burst (glitches)

- Caused by instruments or environment (known or unknown)
- Diminish scientific data available
- Hinder GW detection (mask and/or mimic)

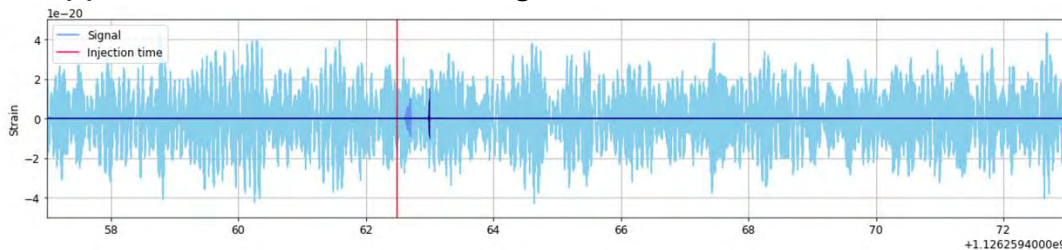


Example of a blip glitch (left) and a intermediate-mass black hole (right)

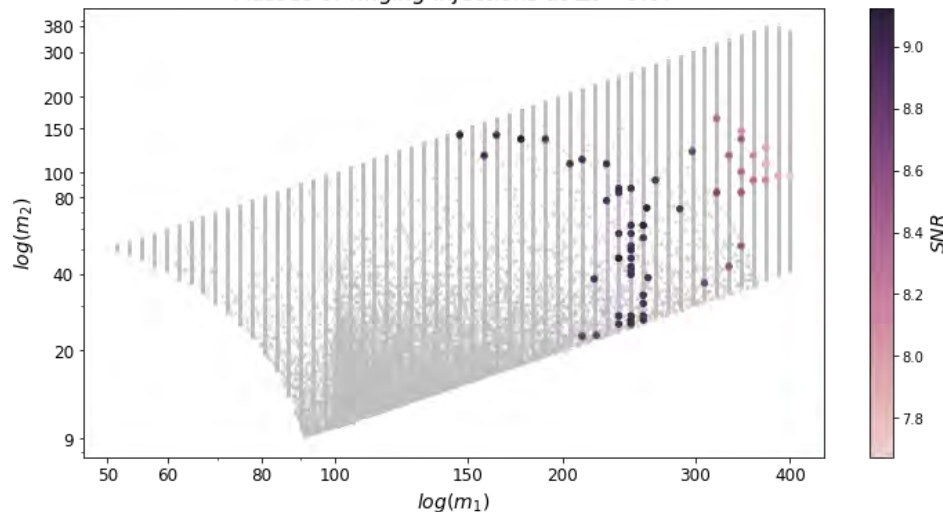


# A simulated GW through a detection pipeline

$\Delta t$ : time when trigger happened – time when GW signal was added to the noise

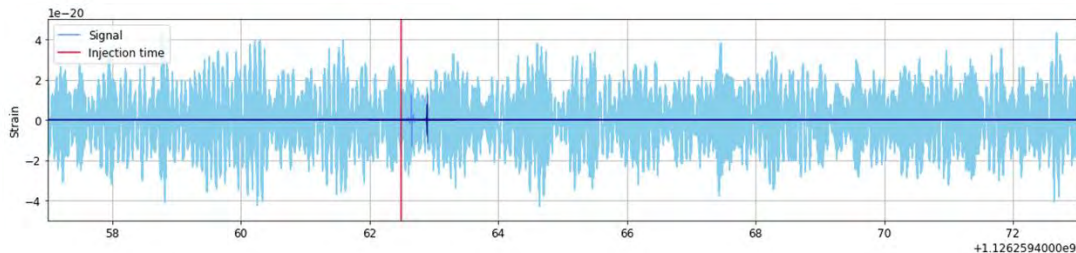


Masses of ringing injections at  $\Delta t = 0.07$

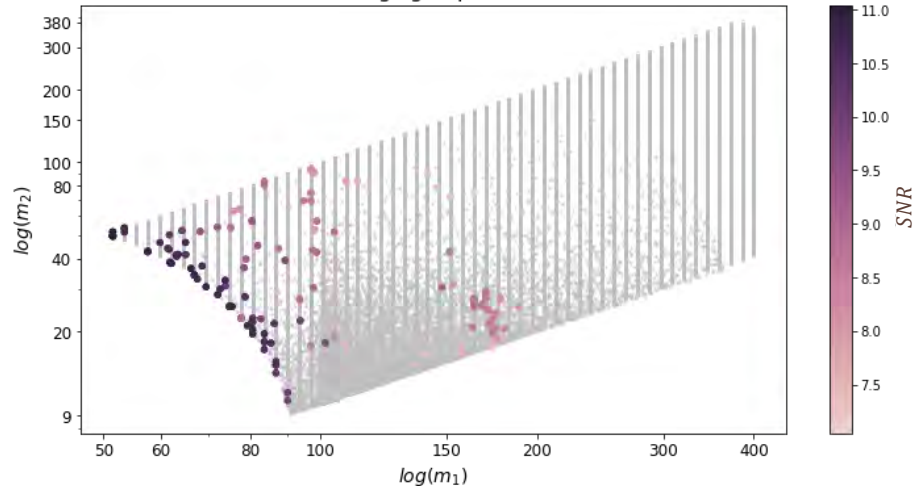


# A glitch through a detection pipeline

$\Delta t$ : time when trigger happened – time when glitch happened



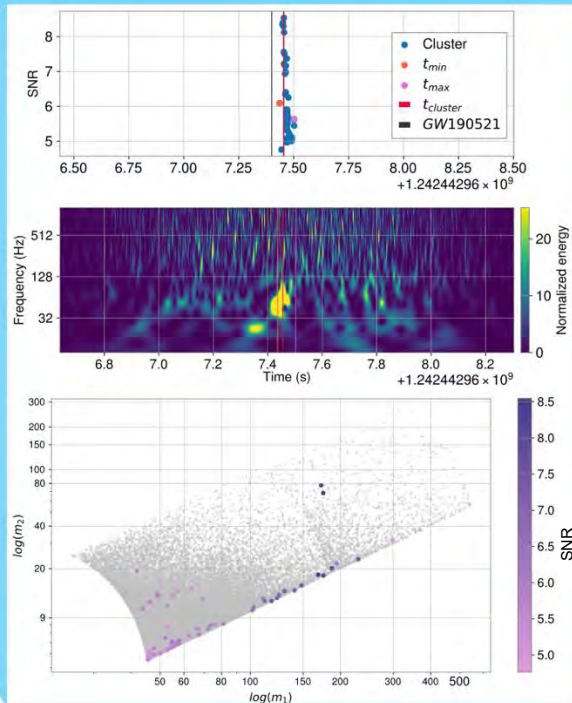
Masses of ringing Blip at  $\Delta t = 0.0455$



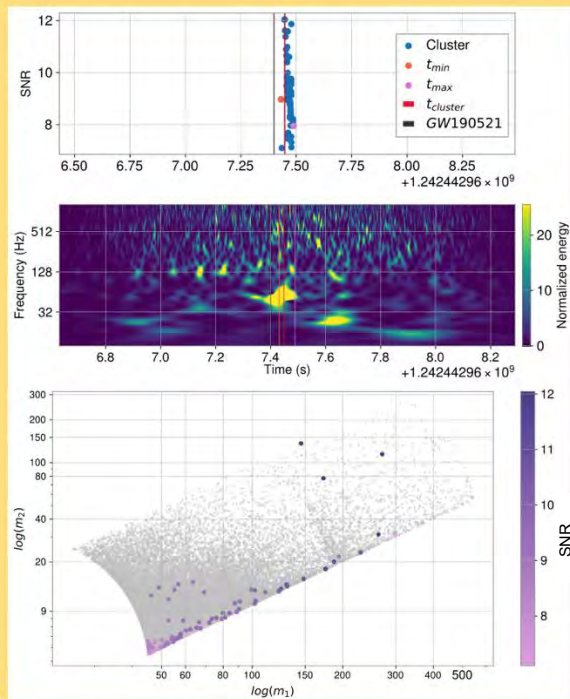


# GW190521

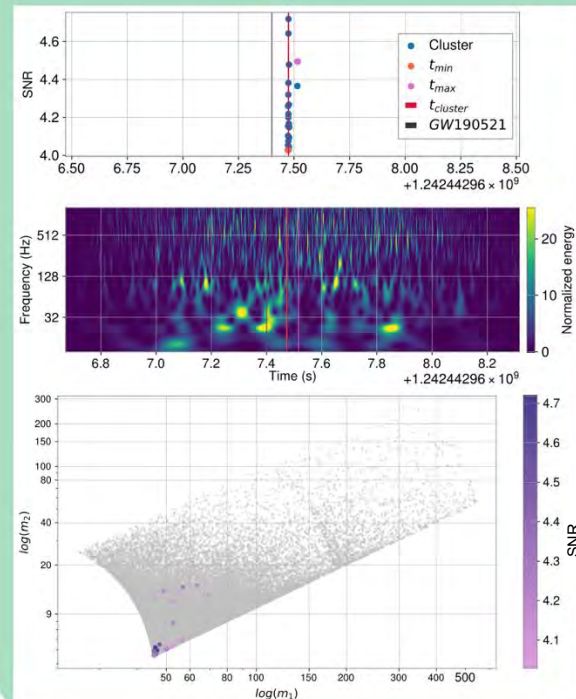
## LIGO Hanford (H1)



## LIGO Livingston (L1)



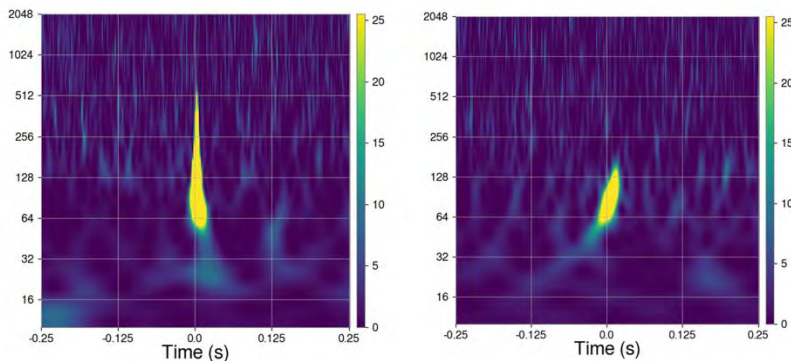
## Virgo (V1)



# Motivation

**Context:** intermediate-mass black holes (IMBH) are the missing link between stellar black holes and supermassive black holes, but they are hard to detect!

**Idea:** use triggers from matched filtering (free information) from detection algorithms to learn the background (glitches) and foreground (GW signals) with ML



Example of a blip glitch (left) and a IMBH (right)

- MF searches use *strict* conditions for detection.
- Can we *relax* the search with the interpolation ability of ML?

# Multi-class classification demo

Demo data from PhysRevD.111.103020 - arXiv 2412.17169

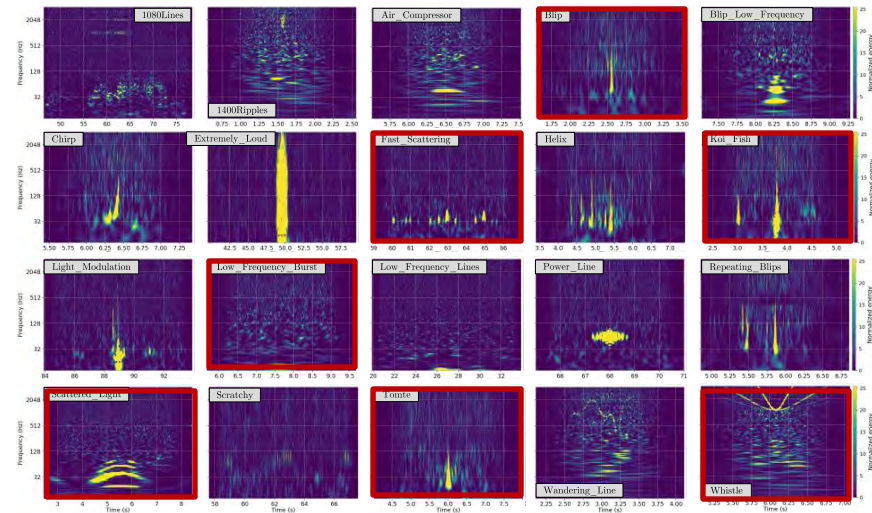
**Task:** Distinguish IMBH from different glitch classes in single detector → we have 3 detectors!

**Algorithm:** Multi-layer perceptron (MLP)

**Input:** Adding time is hard, so let's simplify the problem. Each template is defined by  $m_1, m_2, s_{1z}, s_{2z}, \chi^2, SNR$ . We weight average by SNR to get the feature vector

$$\mu(m_1, m_2, s_{1z}, s_{2z}, \chi^2, SNR)$$

**Output:** class probability

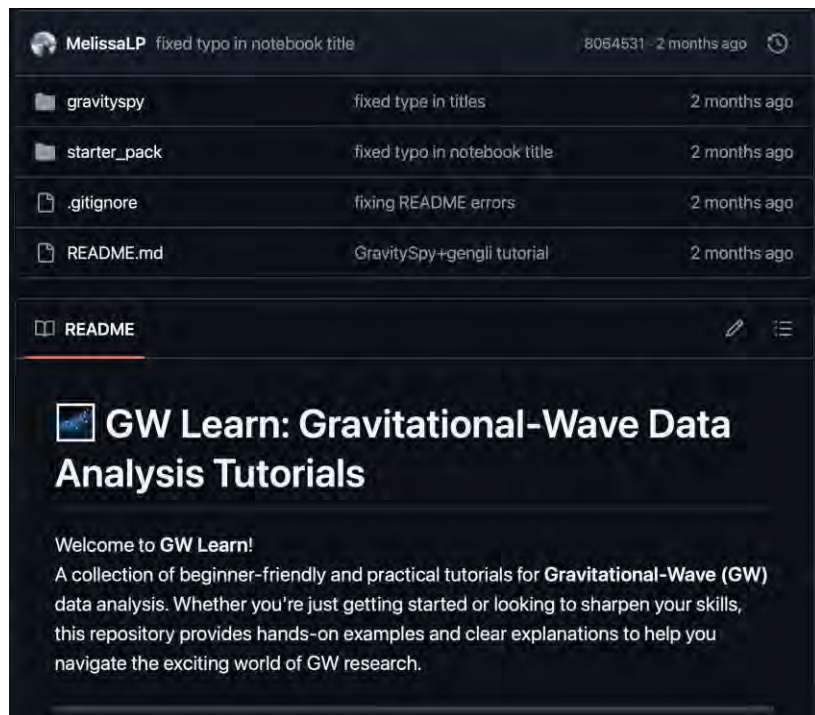


**Idea:** MLP differentiates **6 classes**: 5 different types of background (glitches) and single foreground (GW signals). It uses only **6 parameters** in **single detector**



# About today's tutorial

⚠ Promotion time! [gw\\_learn tutorials](#)



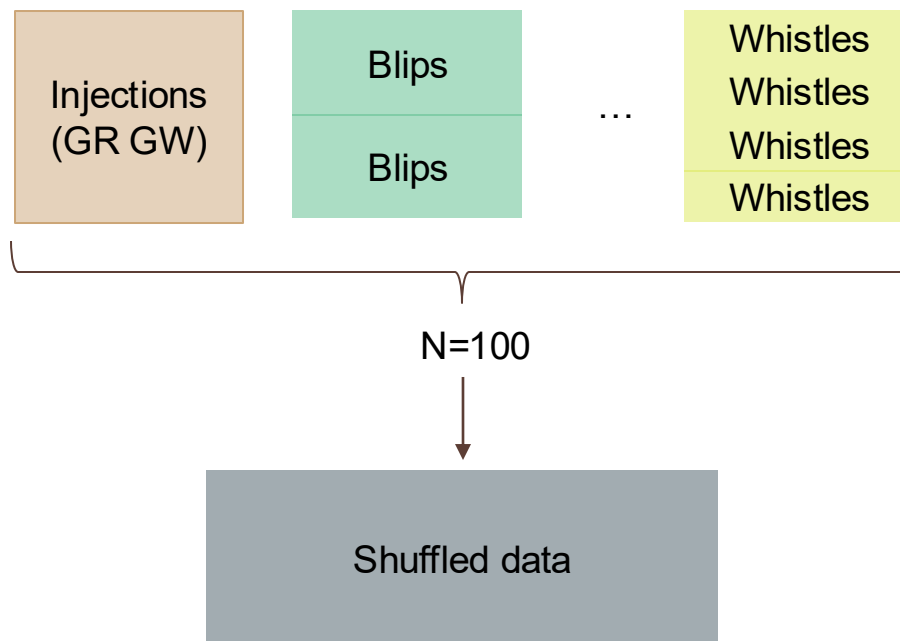
The screenshot shows a GitHub repository interface. At the top, the user 'MelissaLP' is listed with a commit message 'fixed typo in notebook title' and a timestamp '8064531 · 2 months ago'. Below this is a list of files: 'gravityspy' (fixed type in titles, 2 months ago), 'starter\_pack' (fixed typo in notebook title, 2 months ago), '.gitignore' (fixing README errors, 2 months ago), and 'README.md' (GravitySpy+gengli tutorial, 2 months ago). The 'README' file is selected, showing a title 'GW Learn: Gravitational-Wave Data Analysis Tutorials' with a small icon. The content of the README includes a welcome message: 'Welcome to GW Learn! A collection of beginner-friendly and practical tutorials for Gravitational-Wave (GW) data analysis. Whether you're just getting started or looking to sharpen your skills, this repository provides hands-on examples and clear explanations to help you navigate the exciting world of GW research.'

Access tutorial of today:

<https://shorturl.at/vxMH4>

# Dealing with imbalanced data

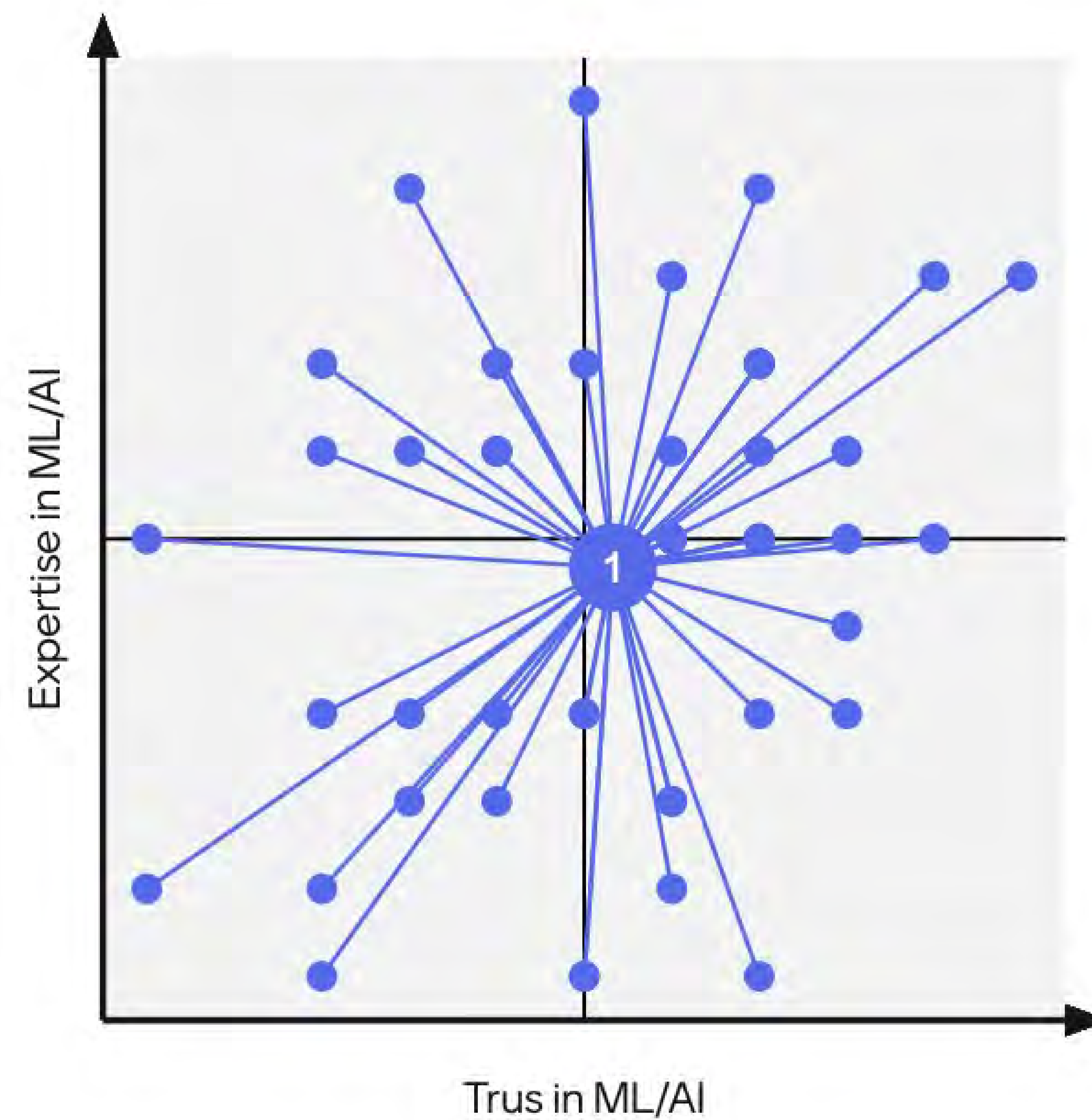
1. Accounting for imbalanced data (bootstrapping with replacement)



Access the tutorial: <https://shorturl.at/vxMH4>



# Where do you stand with ML/AI?



1 Yourself



# Exercise 1: What is bootstrapping and why is it important?

This is a trick question --  
E. T. Jaynes dixit

I'm not sure — I'm a physicist  
and I've ever heard of  
bootstrapping. Is it like  
dimensionality reduction or  
some kind of data compression  
technique?

Resampling

Computing  
uncertainties when you  
don't know what else to  
do

resampling

A way to estimate the unknown  
sampling/data distribution  
using the current data to derive  
estimators for target (usually  
Frequentist) statistics.

When you want to report  
uncertainties but don't have  
the ability/compute/will to draw  
additional samples from the  
distribution of interest.

I think it's important for  
uncertainty  
quantification



# Exercise 1: What is bootstrapping and why is it important?

Getting data is hard -  
how can we maximize  
what we have?

to generate more data  
and reduce potential risk  
of biases

Adding: Bootstrapping cannot  
fix out-of-distribution problems,  
since it assumes the unknown  
target distribution and the  
current distribution are drawn  
from the same underlying  
sampling process.

Generate new data points  
conforming to certain  
fiducial statistical  
properties

to increase ML accuracy  
at lower FARs

Low cost generation of  
samples whose statistical  
properties remain faithful to  
the underlying process one  
is trying to sample from.

When more data is created by  
using previous data. It is useful  
if the data you have is not  
enough for proper analysis, and  
it's not feasible to get more  
data using standard methods.

Bootstrapping is a statistical  
resampling method where  
samples are drawn with  
replacement from the original  
dataset. It's important because  
it allows us to estimate the  
uncertainty of a model.  
ChatGPT

Quote

– *Who*

## Exercise 2: What are the main problems with this network?

Too shallow

Not wide enough

too small.

Data are not normalized

It's very small — it only has two layers. Maybe include more layers and also include a variable learning rate, as well as dropout. Maybe it's overfitting? I noticed the accuracy goes up only to .75.

Input data not scaled appropriately

Neural network architecture is a "prior" over the structure you expect in your data. MLP assumes just a general curve, and with few parameters/layers the functions are probably too simple.

Change the activation to RELU may better performance.



## Exercise 2: What are the main problems with this network?

6x more glitch examples than signal examples

relu better for training than tanh

Significant confusion between injections and low frequency burst. Probably needs to lower frequency high pass while simulating injection samples?

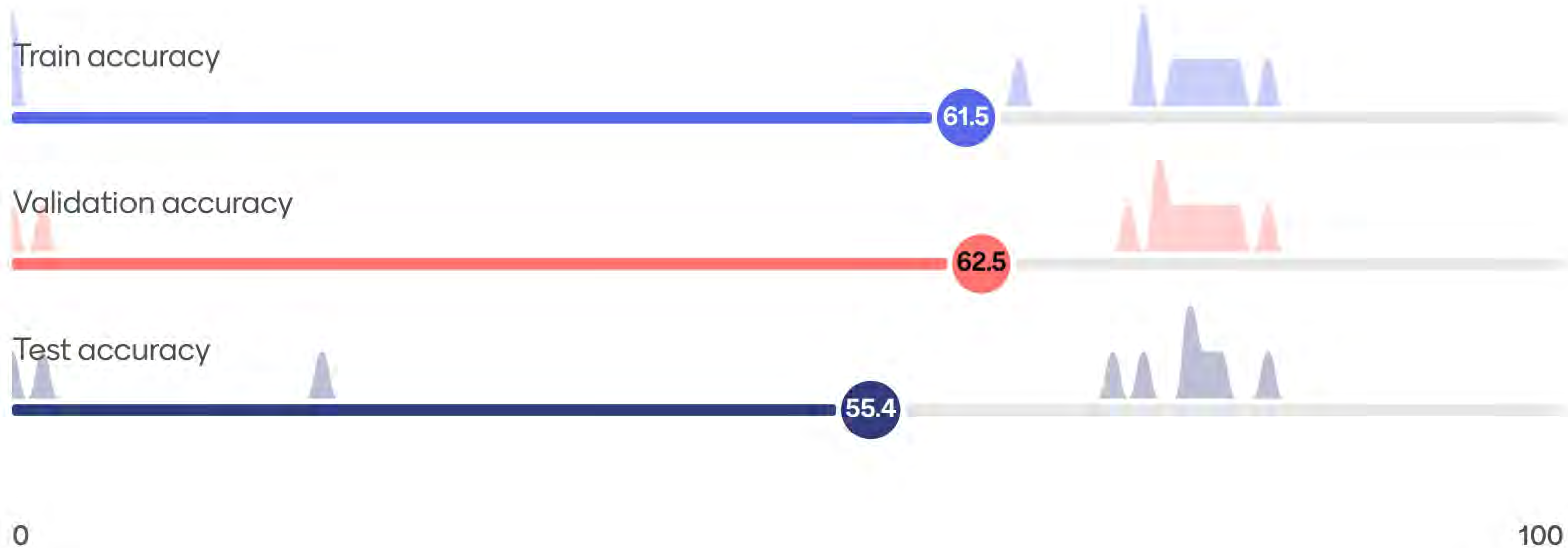
Swish is even better than ReLU

maybe its small

Just keep running can improve

ML is trial and error

## Exercise 3: How much did your network improve?





# This is the end of the tutorial! What did you learn?

great tutorial the architecture for mlps  
ml is trial and error mlp  
swish selu keras architecture is imp  
tuning ml is painful useful tricks  
hourglass architecture  
work with many classes applied ml in gw  
activation functions

D1: Sofia Alvarez-  
Lopez

S1: Rodrigo Tenorio

D2: Mervyn Chan

S2: Andrew Miller

D3: Sidd Soni

S3: Collin Capano

D4: Chayan  
Chatterjee

S4: John Veitch

Access to the solutions:  
<https://shorturl.at/OVfMw>

Senior group chairs

# What would you like to talk about? Short link: <https://shorturl.at/Blil6>

Why do we need ML in GW physics

Hamiltonian dynamics of binary black hole systems!

What is the best data representation for doing ML detchar

How to make GW searches interpretable? (aka how do you make it define a consistent FAR)

Is uncertainty quantification necessary or relevant for ML models used in GW physics?

potential solutions for glitch removal in 3G detectors' era

Critically evaluating where ML can uniquely improve GW detection vs. improved implementations/understanding of 'classical' techniques

Are there ways to get around the Black Box issue?



# What would you like to talk about? Short link: <https://shorturl.at/Blil6>

ML in the context of waveform modelling

Neural ODEs and PINNs in GWs?

What does detection mean for overlapping signals?

Relationship between glitch identifications and PE.  
Downstream implication (astrophysics / test GR / etc) of erroneous glitch identification / data cleaning

How do we trust ML outputs for detector applications?

ML workflow for overlapping signals

Ways to understand what exactly is happening in the layers of a neural net to learn about novel hidden structures in the data

How can we do background estimation in a foreground of tons of long-lived GW signals?

# What would you like to talk about? Short link: <https://shorturl.at/Blil6>

Matched filter is the "optimal statistic". Can we quantify how much more we can (theoretically) learn by employing ML methods?

using ML for unmodeled searches or parts of parameter space where we don't have templates in traditional searches

Do we need global fit for next 3G GW detection ?

Improved CBC classification methods

a complete ml pipeline vs a traditional pipeline with ml replacing one step in the middle. Is there a balance to be achieved?

Detecting and doing PE of events which are contaminated with glitches using ML

ML in the context of waveform modelling

Do we actually need to remove glitches in 3G (signal very long, who cares if there's a small glitch on top)



# What would you like to talk about? Short link: <https://shorturl.at/Blil6>

how much extra volume improvement can ML bring to the table? Is there a particular region in the search parameter space where ML has a unique advantage?

How do we trust ML outputs for detchar applications?

PINNs in GW

Ways to improve ml training for detchar purposes.

Application of Transformer in GW search

Ways to improve ml training for detchar purposes.

can ML do unmodeled GW searches?

How do we know whether ML is giving us the correct waveform or not for very complicated systems like BBHs in eccentric precessing orbits?

# What would you like to talk about? Short link: <https://shorturl.at/Blil6>

Incorporating diagnostic data channels from GW detectors in searches directly by assimilating them using ML stages, and using that to discriminate signals from noise transients invariantly

How to design a search pipeline around ML in practice?

Calibration uncertainty reduction using ML

At which SNR levels will template-banks-based methods become too computationally expensive? Will they at all?

How would we actually verify a detection of an unmodeled signal by a ML algorithm?

What can traditional searches learn from ML?

Is an end to end joint model for noise transients and astrophysical sources feasible with ML?

how often does the distribution of noise change during the run roughly? Or how often do ML models need to be trained/recalibrated during the obs run?

What would you like to talk about? Short link: <https://shorturl.at/Blil6>

ML for  
unmodelled/unknown  
signals