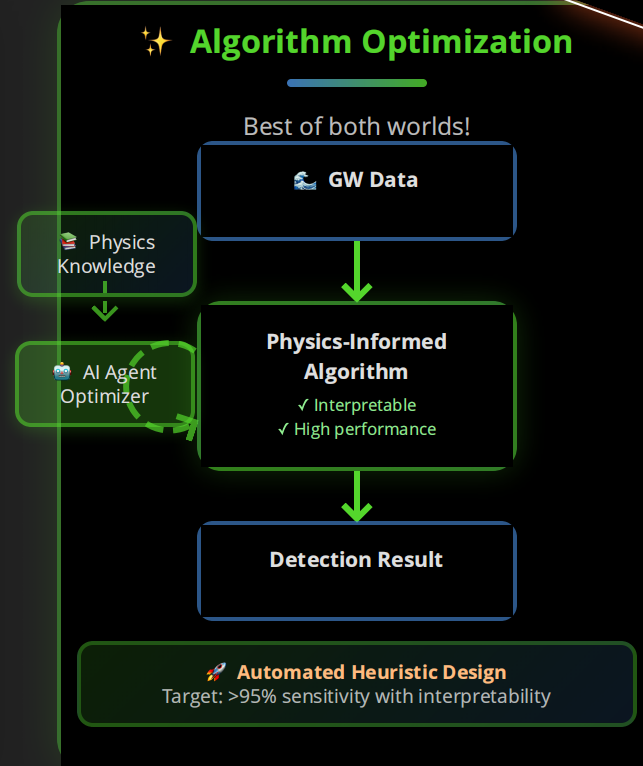
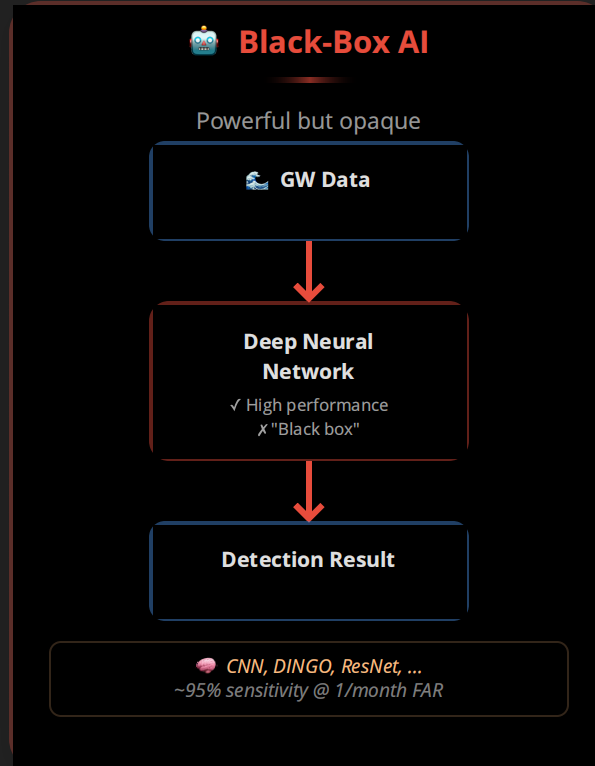
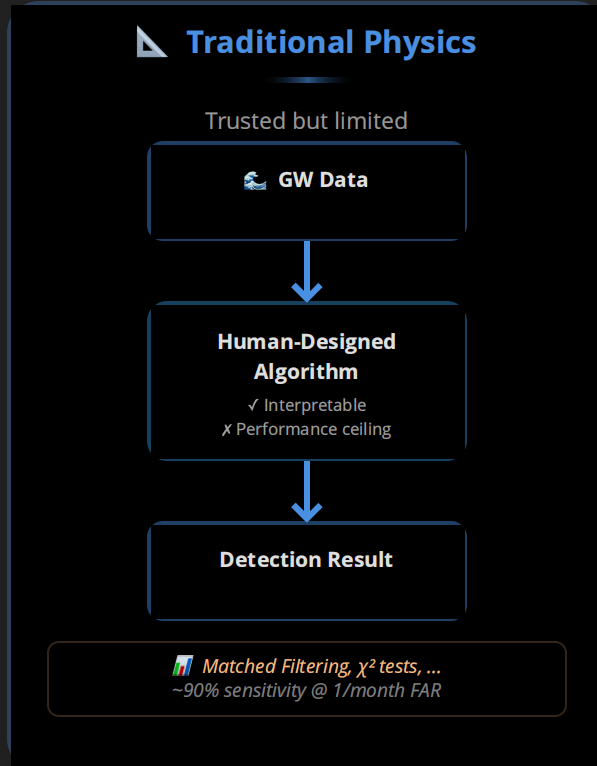


# Interpretable AI for Gravitational Wave Detection

When black-box AI meets physics: Can we have both trust and performance?

OUR WORK

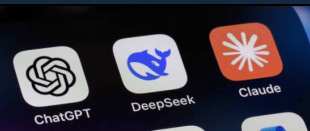


**Key Innovation:** Agent-based LLMs iteratively optimize physics-based algorithms, creating transparent AI systems that scientists can trust and understand

Presenter: He Wang

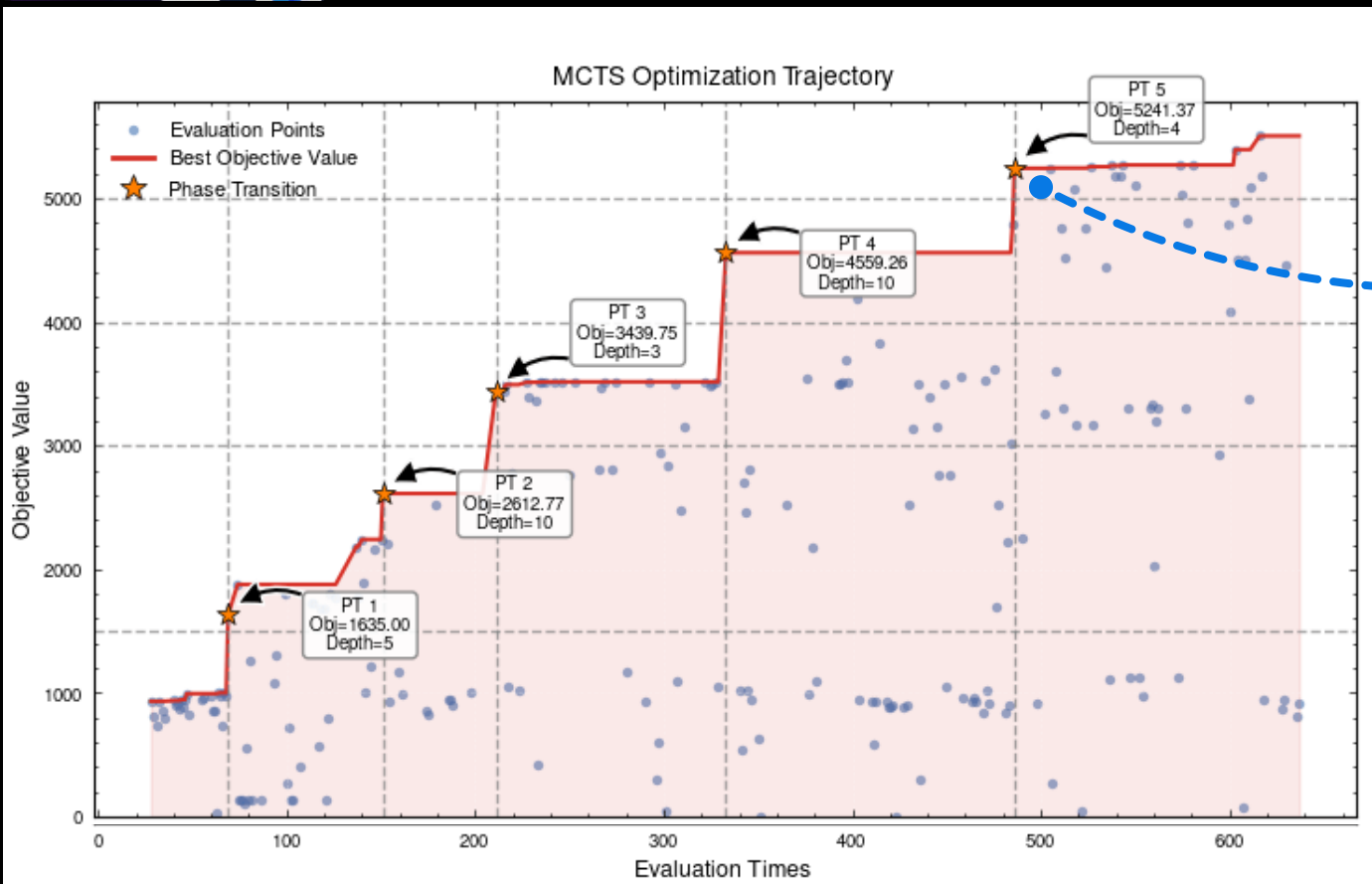
# Agent-Based Optimization Framework for GW Detection

Iterative algorithm development guided by physics-informed AI agents



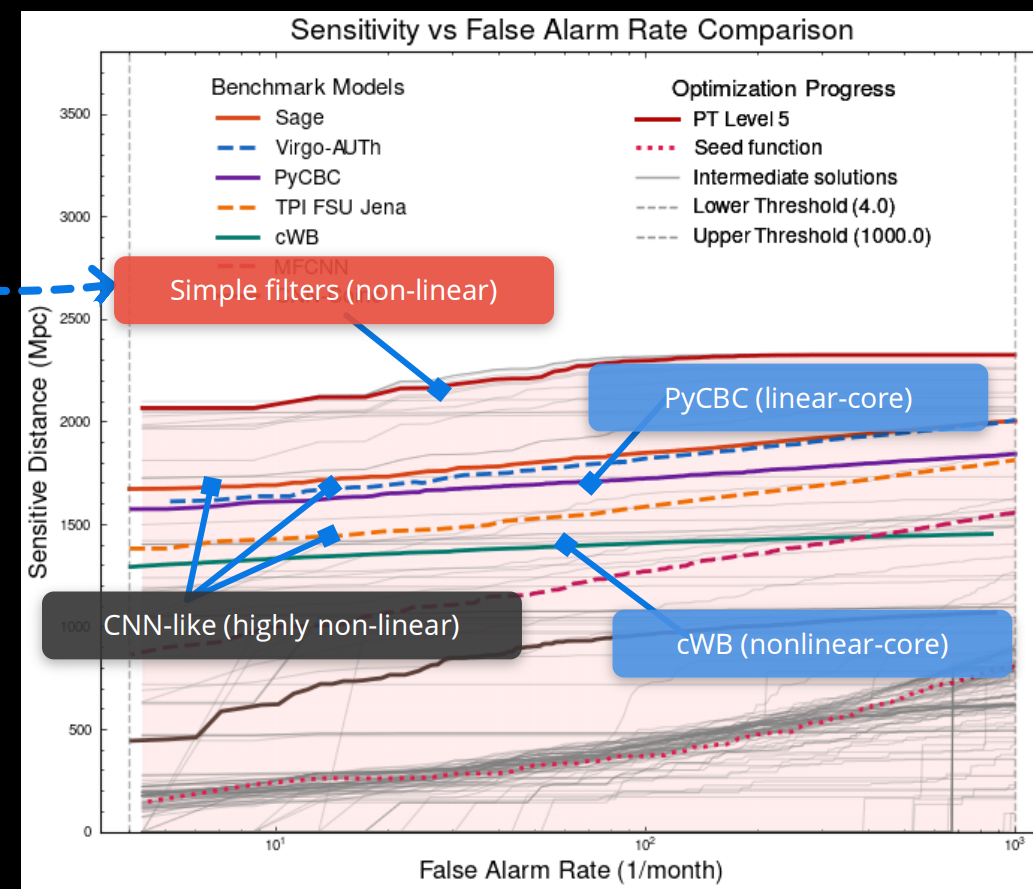
## MCTS Optimization Trajectory

Automated exploration of algorithm parameter space



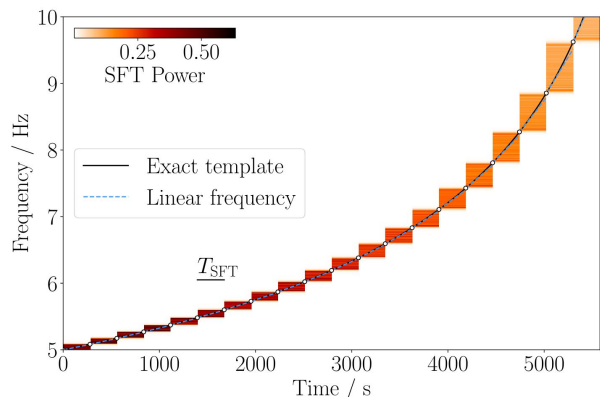
## Sensitivity vs False Alarm Rate Comparison

Benchmarking against state-of-the-art methods



**Framework Advantage:** Agent-based LLMs navigate complex optimization landscapes, achieving **targeted performance improvements** while maintaining **algorithmic interpretability**

# How to analyze long-duration CBC signals with a time-varying detector response, non-stationary noise, and data gaps



$$([d, h]) \doteq \Delta f \sum_{\alpha=0}^{N_{\text{SFT}}-1} A_{\alpha} \Lambda_{\alpha} e^{i\varphi_{\alpha}} \mathcal{C}(f_{\alpha}, \dot{f}_{\alpha}; \tilde{d}_{\alpha})$$

**Use Short Fourier Transforms!**

**Stop** using this!

$$\langle x, h \rangle = 4 \operatorname{Re} \int_0^{\infty} df \frac{\tilde{x}^*(f) \tilde{h}(f)}{S_n(f)}$$



Long-duration signals are hard to analyze in frequency domain.

“2G methods but longer”-approach won’t cut it.

SFTs provide (100–1000)x acceleration and data compression and solve all problems related to long durations.

Rodrigo Tenorio  
rodrigo.tenorio@unimib.it  
U. Milano-Bicocca

Tenorio & Gerosa  
PRD 111, 104044 (2025)

[github.com/rodrigo-tenorio/sfts](https://github.com/rodrigo-tenorio/sfts)

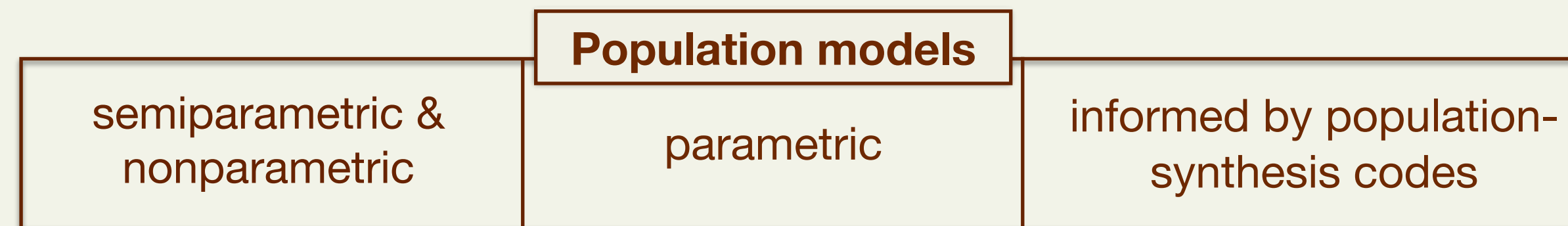
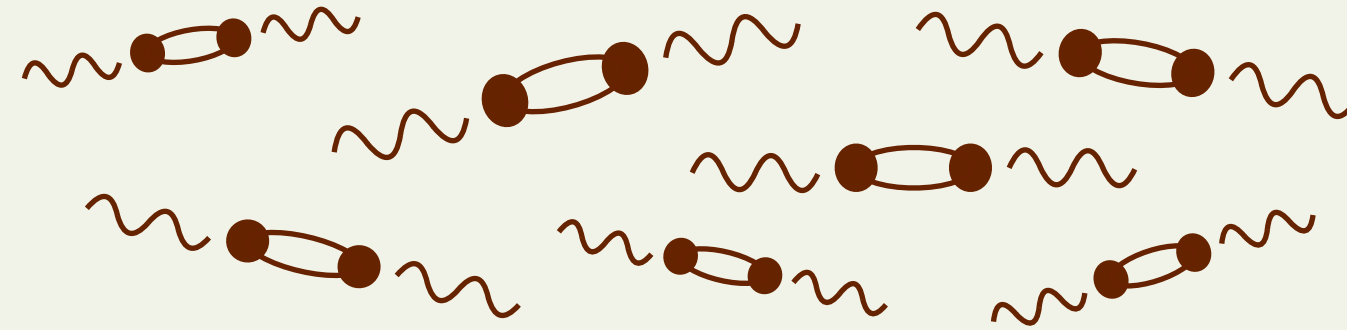


# Reconstructing parametric gravitational-wave population fits from nonparametric results without refitting the data

**C.M. Fabbri** <sup>(1, 2)</sup>, **D. Gerosa** <sup>(1, 3)</sup>, **A. Santini** <sup>(4)</sup>, **M. Mould** <sup>(5, 6)</sup>, **A. Toubiana** <sup>(1, 3)</sup>, **J. Gair** <sup>(4)</sup>

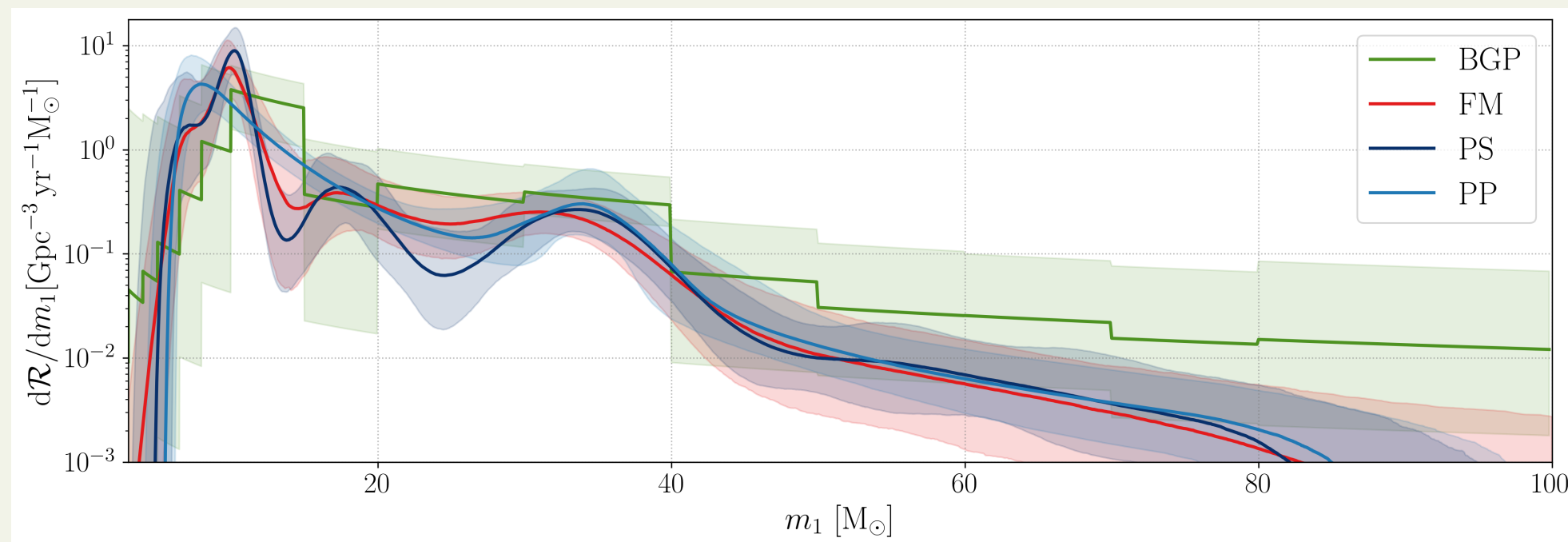
Phys. Rev. D 111 (2025) 104053,  
arXiv:2501.17233

# Developing new approaches for population fits



## Current approach:

Hierarchical Bayesian inference



Fit the dataset multiple times assuming different population models

## Our approach:

Can we obtain multiple population fits analyzing data once?

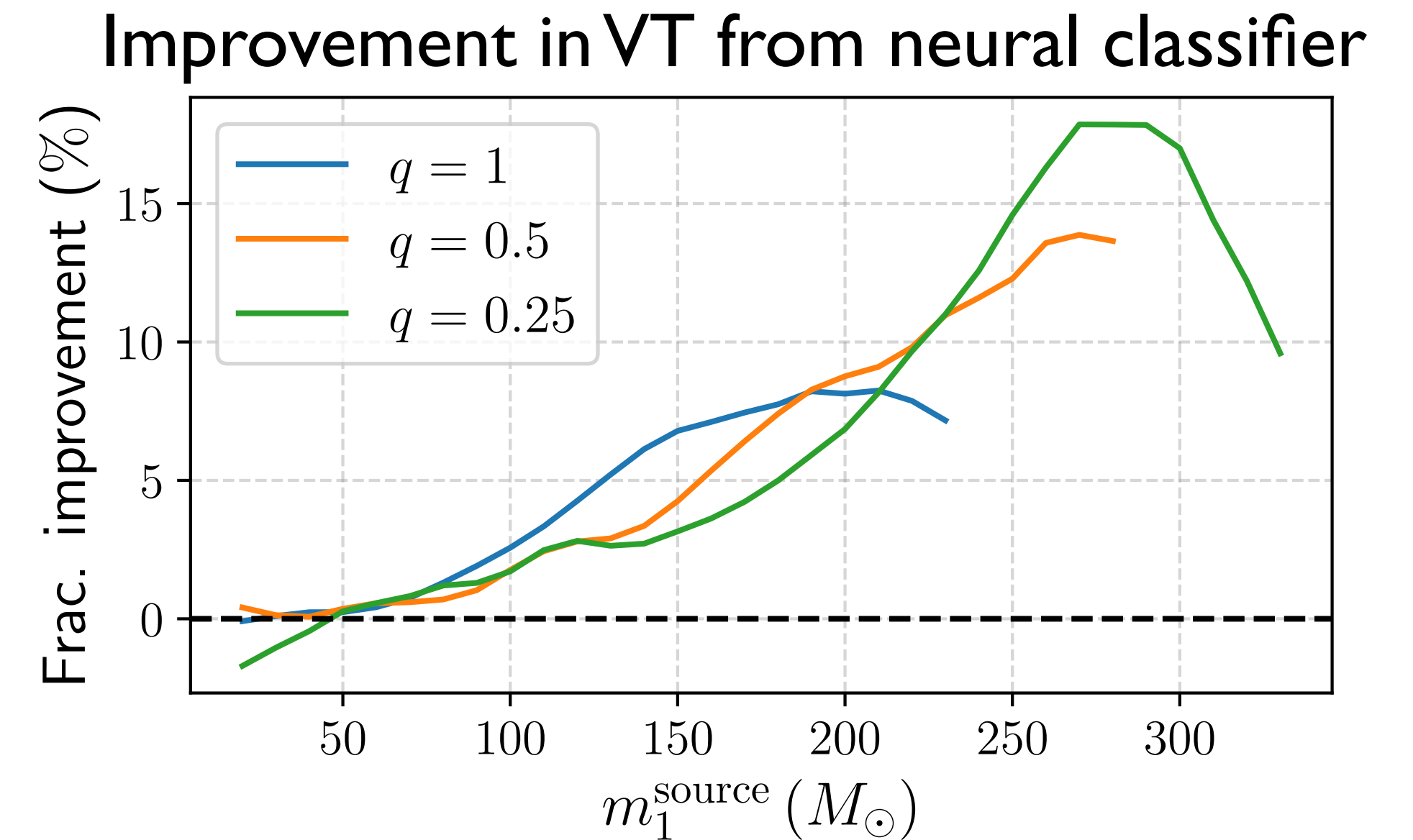


Can we get  $p(\lambda | d)$  without refitting data?

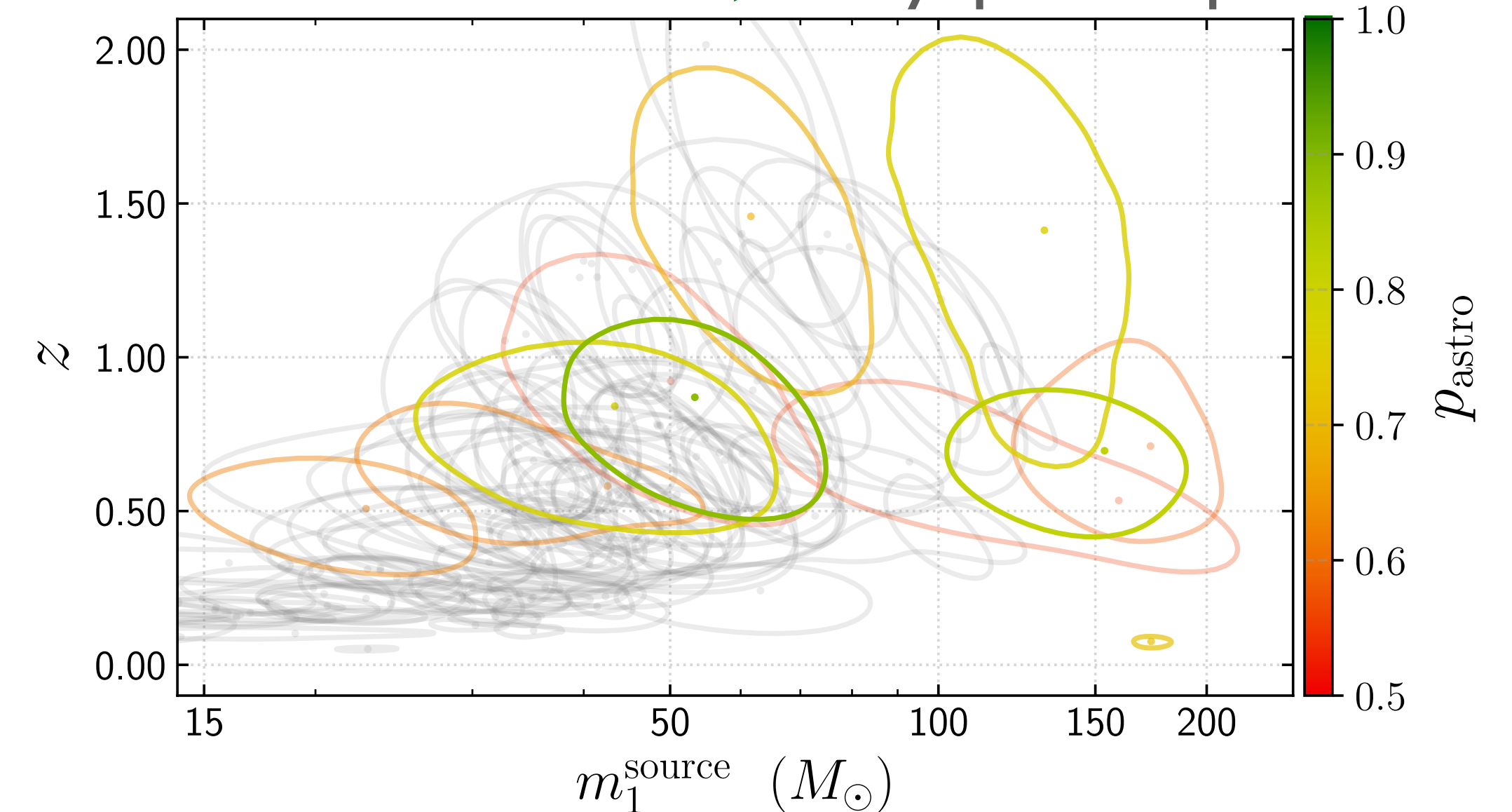
# Best of both worlds: combining principled gravitational wave searches with AI/ML

*(Jay) Digvijay Wadekar (JHU)*

- Combining ML methods with principled matched-filtering based searches can help:
  1. build leaner ML tools which are more robust (w.r.t generalizability, interpretability)
  2. improve upon simplistic analytic assumptions (e.g., noise is Gaussian)
- We combine the IAS search pipeline with
  1. **Autoencoders / decision trees** to build more efficient template banks
  2. **Neural classifiers** for down-weighting noise transients (increases sensitivity volume)
  3. **Normalizing flows** for efficient marginalization over higher harmonics (and precession, eccentricity) d.o.f



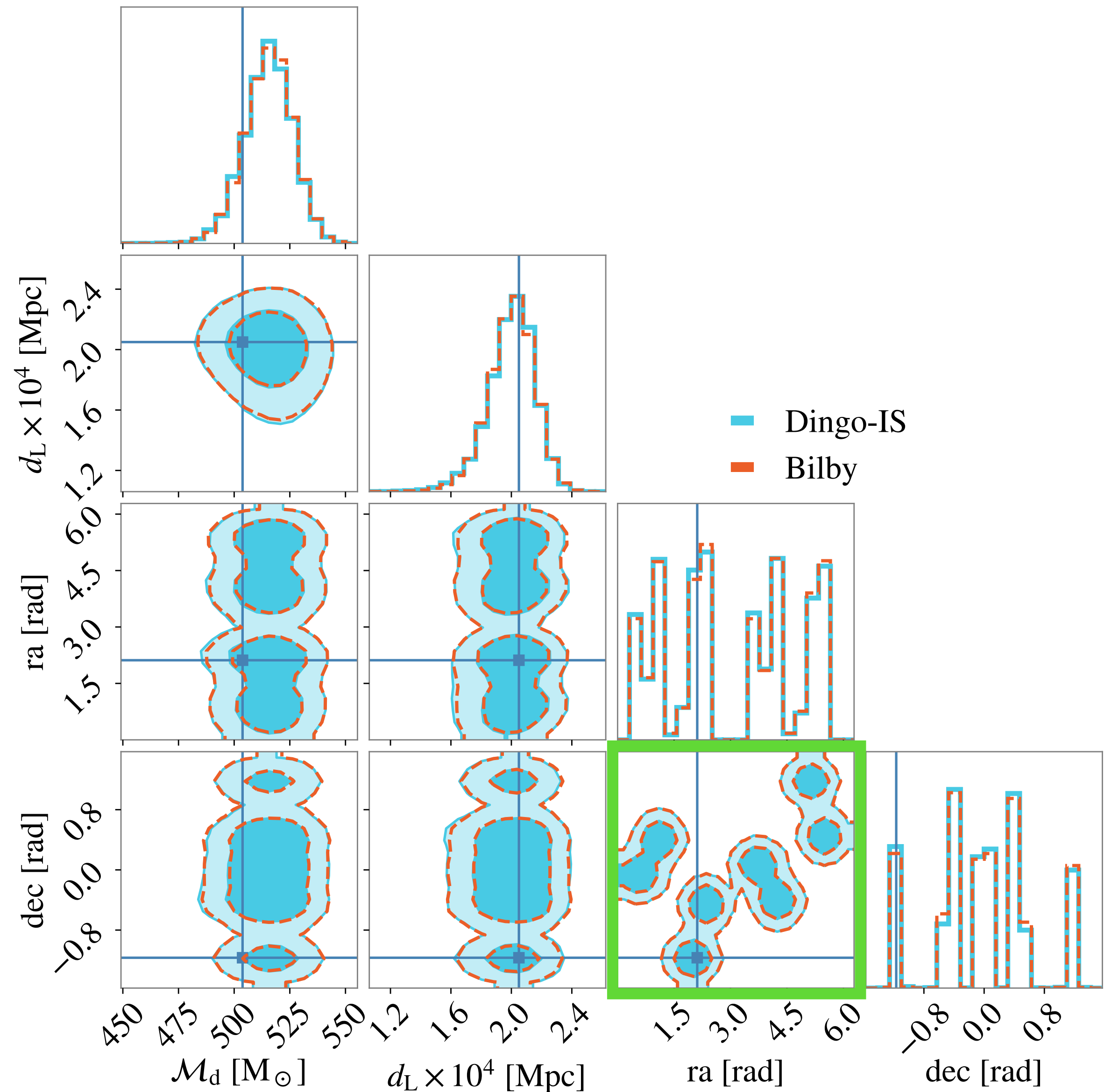
Colored: New detections, Gray: prev. reported



# Fast and accurate parameter estimation of high-redshift sources with the Einstein Telescope

**Filippo Santoliquido,**

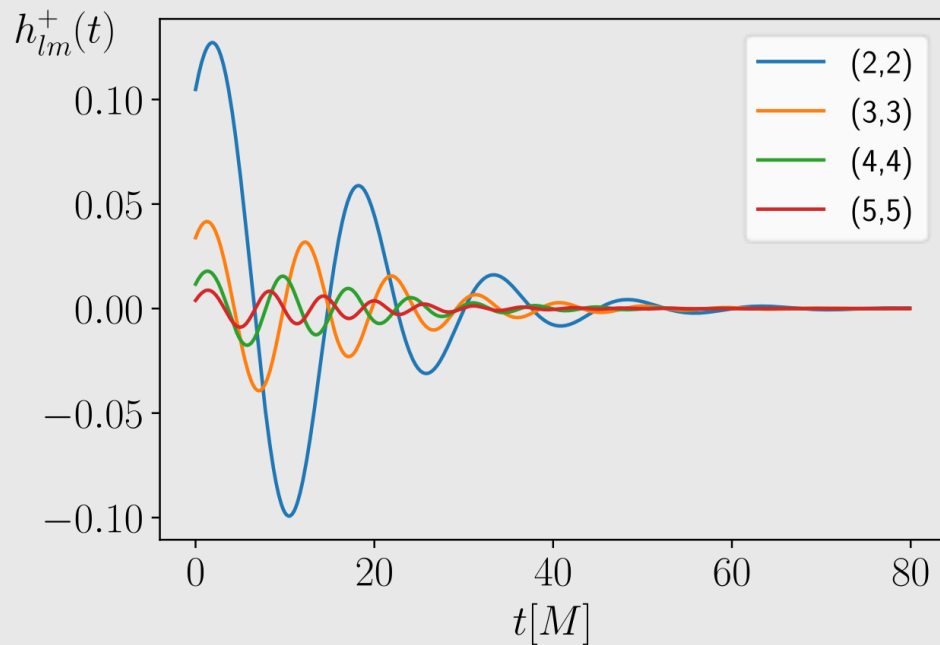
Jacopo Tissino,  
Ulyana Dupletsa,  
Jan Harms,  
Marica Branchesi,  
Manuel Arca Sedda,  
Maximilian Dax,  
Annalena Kofler,  
Stephen R. Green,  
Nihar Gupte,  
Isobel M. Romero-Shaw,  
Emanuele Berti



# RINGDOWN MODE AMPLITUDES OF PRECESSING BINARY BLACK HOLES

Francesco Nobili, S. Bhagwat, C. Pacilio, D. Gerosa

✉ fnobili@uninsubria.it



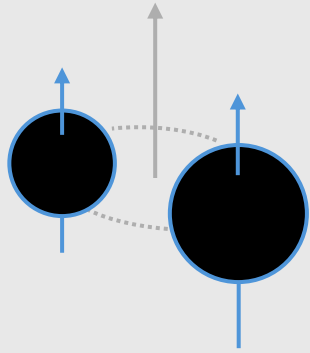
- Mode amplitudes depend on **merger** phase

- Modeling amps

Tests of GR

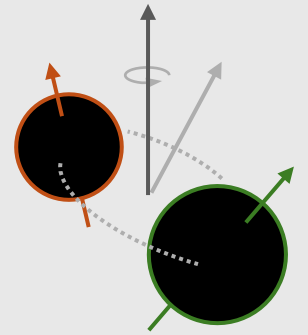
RD parameter estimation

## SPIN-ALIGNED



- **Simpler:** 3D parameter space
- Solid phenom description
- Several models

## PREPROCESSING



- **Complex:** 7D parameter space
- Only 1 paper on phenom
- **no models!**

## OUR WORK

Amp **correlations** with  
BBH physical parameters

**models** using  
Gaussian Process Regression